

Swift Observation of GRB 070724A

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0 Revisions

Ground observations of candidate host galaxy and other sources in the XRT error circle are updated. The XRT light curve and slopes are update and the break time added.

1 Introduction

BAT triggered on GRB 070724A at 10:53:50 UT (Trigger 285948) (Ziaeepour, et al., *GCN Circ.* 6654). This was a 0.512 sec rate trigger with 13.03 significance on a short burst with $T_{90} = 0.4 \pm 0.04$ sec (Parsons, et al., *GCN Circ.* 6656). Swift slewed to this burst immediately and XRT began follow-up observations at $T + 72.1$ sec, and UVOT at $T + 75$ sec. Our best position is the XRT location $RA(J2000) = 27.80815$ deg (01h51m13.96s), $Dec(J2000) = -18.59448$ deg ($-18d35'40.1''$) with an error of 2.2 arcsec (Page, et al. *GCN Circ.* 6659). No optical/UV counterpart was found for this burst up to a 3σ magnitude limit of 20.4 in White filter (160 – 650 nm) at $\sim T + 100$ sec (De Pasquale, et al., *GCN Circ.* 6660). In the Digital Sky Survey (DSS) images it was found that a faint blue source partially overlaps the error circle of the refined XRT position (Bloom, *GCN Circ.* 6658, Bloom & Butler, *GCN Circ.* 6661). The same source has been detect in some of the UVOT filters. An IR observation of this burst by UKIRT (Levan, et al., *GCN Circ.* 6662), P60, Gemini (Cenko, et al., *GCN Circ.* 6664, Cucchiara, et al., *GCN Circ.* 6665), and VLT (Covino, et al., *GCN Circ.* 6666) confirms the presence of this source, its classification as a galaxy and its redshift $z = 0.457$. In the second observation by Gemini/GMOS (Cenko, et al., *GCN Circ.* 6664) a faint source of magnitude ~ 24.6 in i' has been found at $\sim T + 2.38$ hours. The position of this source is $RA(J2000) = 01h51m14.05s$, $Dec(J2000) = -18d35'42.1''$. Further observations of this source by the VLT (D'Avanzo & S. Covino, *GCN Circ.* 6667) show that it does not vary. The VLA observation of GRB 070724A finds 4 sources in the XRT error circle, including both the candidate host galaxy and the faint source reported by Cenko, et al., *GCN Circ.* 6664, D'Avanzo & S. Covino, *GCN Circ.* 6667.

2 BAT Observation and Analysis

Using the data set from $T - 310$ to $T + 310$ sec, further analysis of BAT GRB 070724A has been performed by Swift team (Parsons, et al., *GCN Circ.* 6656). The BAT ground-calculated position is $RA(J2000) = 27.824$ deg (01h51m17.9s), $Dec(J2000) = -18.610$ deg ($-18d36'35''$) ± 1.2 arcmin, (radius, systematic and statistical, 90% containment). The partial coding was 91% (the offset angle was 16.44 deg).

In the masked-weighted light curves (Fig.1) a peak with substructures starts at $\sim T - 0.1$ sec and returns to background at about $\sim T + 0.3$ sec. The maximum is at ~ 0.2 sec. No extended emission was observed later. T_{90} (15 – 350 keV) is 0.4 ± 0.04 sec (estimated error including systematics).

The time-averaged spectrum from $T - 0.0$ to $T + 0.4$ sec is best fitted by a simple power law model. This fit gives a photon index of 1.81 ± 0.33 , ($\chi^2 = 51.99$ for 57 d.o.f.). For this model the total fluence in the 15 – 150 keV band is $(3.0 \pm 0.7) \times 10^{-8}$ ergs cm^{-2} and the 1-sec peak flux measured from $T - 0.3$ sec in the 15 – 150 keV band is 1.0 ± 0.2 ph $\text{cm}^{-2} \text{sec}^{-1}$. All the quoted errors are at the 90% confidence level.

3 XRT Observations and Analysis

The Swift-XRT data obtained for GRB 070724A are used for refined analysis of the X-ray emission of this burst (Page, et al., *GCN Circ.* 6659). Using 399 sec of overlapping XRT PC mode and UVOT V-band data, the XRT-UVOT alignment and matching UVOT field sources to the USNO-B1 catalogue the astrometrically corrected XRT position of GRB 070724A is: RA($J2000$) = 27.80815 deg (01h51m13.96s), Dec($J2000$) = -18.59448 deg (-18d35'40.1") \pm 2.2 arcsec (90% confidence). This position is within 2.7 arcsec of the initial XRT position (Ziaeeppour, et al. *GCN Circ.* 6654) and 77.8 arcsec from the ground-calculated BAT position (Parsons, et al., *GCN Circ.* 6656).

The 0.3 – 10 keV light curve (Fig.2) shows an initial extended emission/flare that arrives to its maximum at around $\sim T+127$ sec, followed by a very steep decline and a flare peaked at $\sim T+200$ sec and fading with a slope of ~ 2.2 . From $\sim T+400$ sec a shallow regime begins with a slope of $0.65^{+0.10}_{-0.12}$. Finally a break was observed at $\sim T+46$ ksec and the slope steepens to 3^{+2}_{-1} .

In order to avoid any spectral evolution during the flares, Photon Counting (PC) mode data from the second and third, $\sim T+4$ to 12 ksec, were used to fit the X-ray spectrum. It can be modelled by a power-law of $\Gamma = 1.3^{+0.8}_{-0.6}$, absorbed by the Galactic column in this direction of $1.43 \times 10^{20} \text{ cm}^{-2}$ (Kalberla, et al., 2005). The average observed flux in 0.3 – 10 keV during this time is $5.18 \times 10^{-13} \text{ ergs cm}^{-2} \text{ sec}^{-1}$ corresponding to an unabsorbed flux of $5.27 \times 10^{-13} \text{ ergs cm}^{-2} \text{ sec}^{-1}$.

4 UVOT Observation and Analysis

The UVOT began observing the field of GRB 070724A at 10 : 35 : 10 UT, 75 sec after the initial BAT trigger (De Pasquale et al., *GCN Circ.* 6660). No optical afterglow is detected by Swift/UVOT in the XRT error circle in the white (98 sec) and V (393 sec) finding exposures, or in the co-added images in any filter down to 3σ magnitude limit. Upper limits are summarized in Table 1. These upper limits are not corrected for Galactic extinction $E(B - V) = 0.01$. The candidate host galaxy (Bloom, *GCN Circ.* 6658, Bloom & Butler, *GCN Circ.* 6661, Levan, et al., *GCN Circ.* 6662, Cenko, et al., *GCN Circ.* 6664, Cucchiara, et al., *GCN Circ.* 6665, Covino, et al., *GCN Circ.* 6666) is also detected by the UVOT in White and UWM2 filters.

Filter	T_{mid} sec	Exposure (sec)	3σ Mag.UL
White	76 – 176	98	20.4
V	182 – 581	393	19.5
White	76 – 5531	380	21.1
V	182 – 5941	806	20.0
B	662 – 5325	216	20.3
U	636 – 6456	332	20.0
UVW1	612 – 6352	432	20.4
UVM2	587 – 6146	432	20.4
UVW2	692 – 5736	432	20.7

Table 1: Magnitude limits from UVOT observations

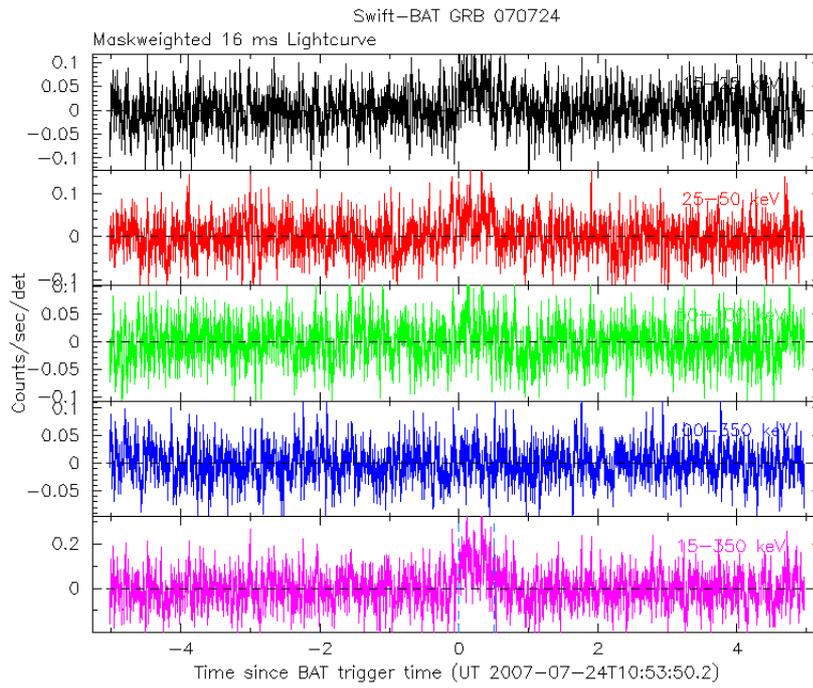


Figure 1: BAT Light curve. The mask-weighted light curve in the 4 individual plus total energy bands. The units are counts/sec/illuminated-detector and T is 10 : 53 : 50.2 UT.

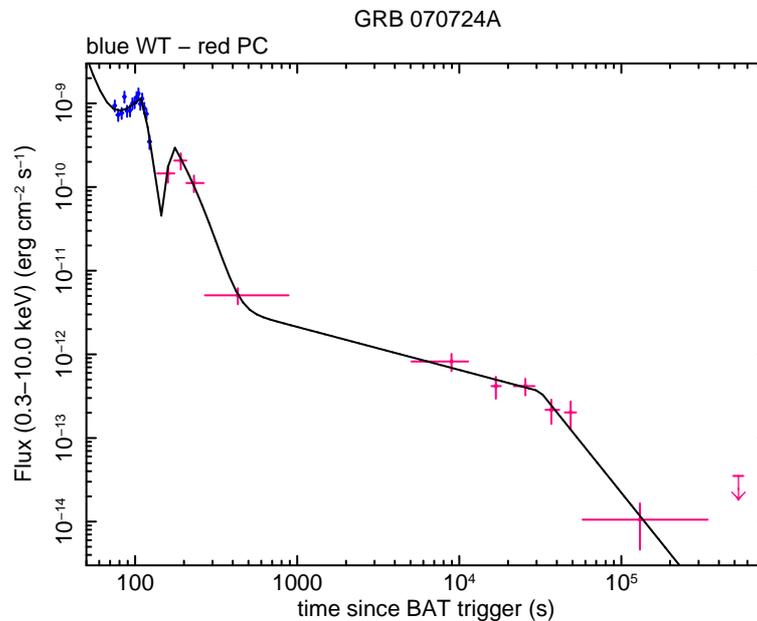


Figure 2: XRT Lightcurve. Flux in the 0.3 – 10 keV band: Window Timing mode (blue), Photon Counting mode (red). The conversion of the absorbed flux is 1 count/sec \iff 5.99×10^{-11} ergs cm $^{-2}$ sec $^{-1}$.